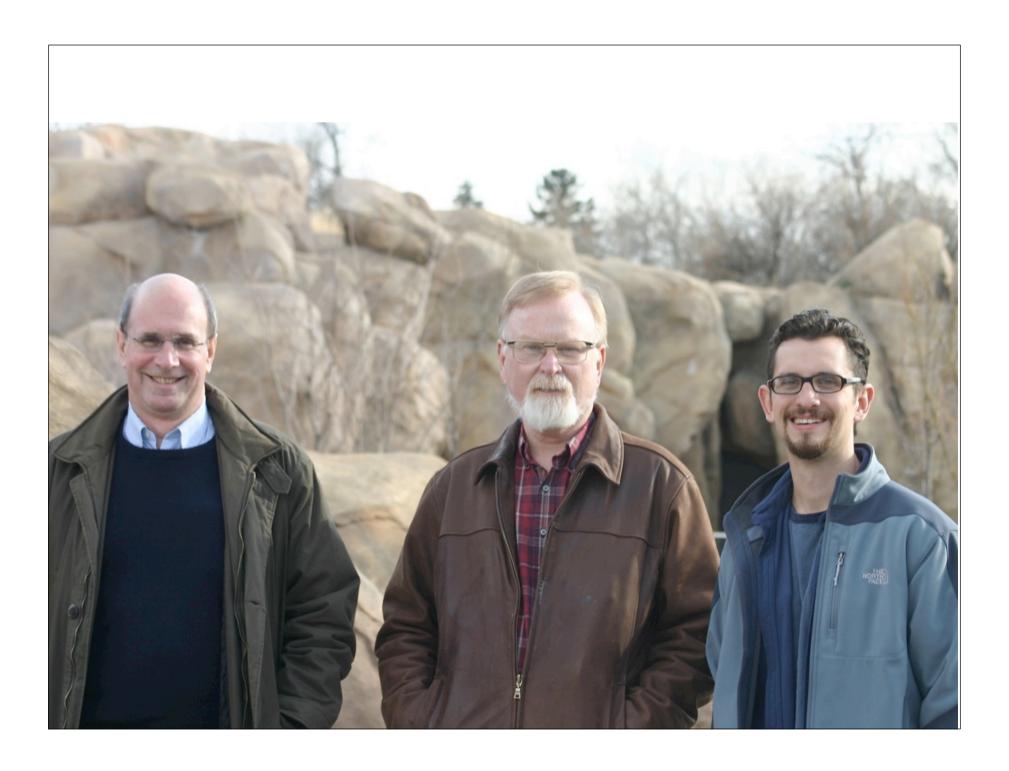
Dynamic Bifurcations and Melting Boundary Convection

Geoff Vasil (CITA, Toronto) Mike Proctor (DAMTP, Cambridge)



MREP60

2010



Consider a general stability problem

$$\frac{du}{dt} = F(u; \alpha)$$

This is the same at the "coupled dynamical system"

$$rac{du}{dt} = F(u; lpha) \qquad rac{dlpha}{dt} = 0$$

What if we make it a bit more interesting?

$$rac{du}{dt} = F(u; lpha) \quad rac{dlpha}{dt} = \epsilon G(lpha)$$

 $\epsilon o 0$ Recovers the previous case

Now it is a (one-way coupled) multi-timescale problem where

$$\alpha = \alpha(\epsilon t)$$

What if we make it even more interesting?

$$rac{du}{dt} = F(u; lpha) rac{dlpha}{dt} = \epsilon G(lpha; u)$$

 $\epsilon o 0$ Recovers the original case

Now it is a (two-way coupled) multi-timescale problem where the bifurcation "parameter" is simply a slowly varying dynamical variable of the system

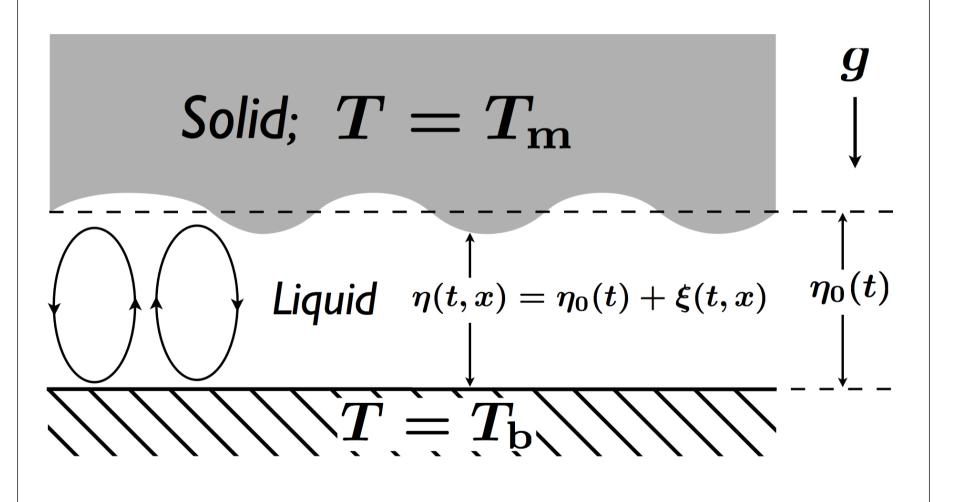
$$rac{du}{dt} = F(u; lpha) rac{dlpha}{dt} = \epsilon G(lpha; u)$$

Rescale time

$$t
ightarrow \epsilon^{-1/2} t, \quad rac{d}{dt}
ightarrow \epsilon^{1/2} rac{d}{dt}$$

$$\epsilon^{1/2}rac{du}{dt}=F(u;lpha)\quad rac{dlpha}{dt}=\epsilon^{1/2}G(lpha;u)$$

A Model Problem



The Stefan Condition

$$S\partial_t \eta + \hat{n} \cdot \nabla T = 0$$
 at $z = \eta(x, y)$

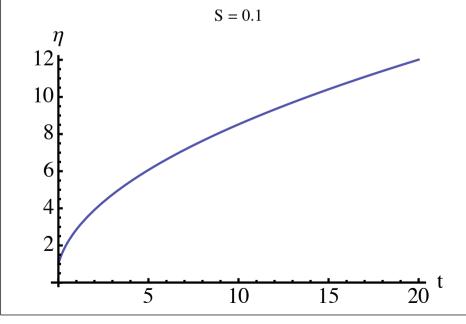
Background State (Similarity Solution)

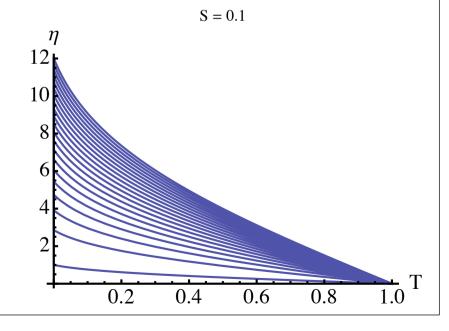
$$\partial_t T_0 = \partial_z^2 T_0 \quad S \partial_t \eta_0 = -\partial_z T_0|_{z=\eta_0}$$

$$rac{1}{S} = \sqrt{\pi} \beta \mathrm{Erf}(\beta) \exp(\beta^2) pprox 2\beta^2 \quad \mathrm{if} \quad S \gg 1$$

$$\eta_0 = \sqrt{4eta^2t+1}$$

$$\eta_0 = \sqrt{4eta^2t+1} \qquad \qquad T_0 = 1 - rac{\mathrm{Erf}(eta z/\eta_0)}{\mathrm{Erf}(eta)} pprox 1 - rac{z}{\eta_0}$$





$$\epsilon = S^{-1/4} \ll 1$$
 Slow time evolution ${
m Ra} = {
m Ra_c} pprox 1708$ $\partial_x = ik_{
m c} + \epsilon\,\partial_X \ \partial_t = \epsilon^2\partial_ au$ $ec u = \epsilon\,A(au,X)ec U_{
m c}(z)e^{ik_{
m c}x} + {
m c.c.}$ $T = 1 - z + \epsilon\,A(au,X)\Theta_{
m c}(z)e^{ik_{
m c}x} + {
m c.c.}$

$$\eta = 1 + \epsilon^2 \tau + \epsilon^3 \Xi(\tau, X) e^{ik_c x} + \text{c.c.}$$

Put it all together...

We get something like the Ginzburg-Landau Equation

$$\partial_{\tau}A = \gamma \tau A + \Xi - |A|^2 A + \partial_X^2 A$$

$$\partial_{ au}\Xi = A$$

$$\gamma = \frac{3 \int_0^1 W_c \Theta_c dz}{|\partial_z \Theta_c|_{z=1}^2} \approx 3.747739$$

These systems have been studied before

Ginzburg-Landau (GL)

$$\partial \tau A = \mu A - |A|^2 A + \partial_X^2 A$$
 $\frac{du}{dt} = F(u; \alpha) \quad \frac{d\alpha}{dt} = 0$

Dynamic GLI

$$\partial_{\tau} A = \gamma \tau A - |A|^2 A + \partial_X^2 A$$

$$\frac{du}{dt} = F(u; \alpha) \quad \frac{d\alpha}{dt} = \epsilon G(\alpha)$$

Dynamic GL2

$$egin{array}{lll} \partial_{ au}A &=& \gamma au A + \Xi - |A|^2 A + \partial_X^2 A \ &\partial_{ au}\Xi &=& A \ rac{du}{dt} &=& F(u;lpha) &rac{dlpha}{dt} &=& \epsilon\,G(lpha;u) \end{array}$$

This system has some interesting new dynamics

Purely Real, No Spatial Dependence

$$\partial_{\boldsymbol{\tau}} A = \mu A - A^3$$
 $A \sim e^{\mu \tau} \text{ for } |A| \ll 1$
 $A \to \pm \sqrt{\mu} \text{ as } \tau \to \infty$

For μ < 0, the solutions decay very rapidly

$$\partial_{ au} A \ = \ \gamma au A - A^3 \quad egin{aligned} A \sim e^{\gamma au^2/2} \ ext{for} \ |A| \ll 1 \ A
ightarrow \pm \sqrt{\gamma au} \ ext{as} \ au
ightarrow \infty \end{aligned}$$

For τ < 0, the solutions also decay very rapidly

Linear, No Spatial Modulation

$$\partial_{ au} A = \gamma au A + \Xi$$
 $\partial_{ au} \Xi = A$

$$A \sim \int_0^\infty \zeta^{rac{1}{\gamma}} \expiggl[-rac{\zeta^2}{2\gamma} \pm \zeta au iggr] \,\mathrm{d}\zeta ext{ for } |A| \ll 1$$

$$A
ightarrow rac{\Gamma\left(1+rac{1}{\gamma}
ight)}{\left| au
ight|^{rac{1}{\gamma}+1}} ext{ for } au \ll 0$$

$$A
ightarrow\sqrt{2\pi\gamma}\left(\gamma au
ight)^{rac{1}{\gamma}}\exp\!\left[rac{\gamma au^2}{2}
ight] ext{ for } au\gg0$$

For τ < 0, the solutions can grow algebraically. There is a morphological instability before there is a convective instability

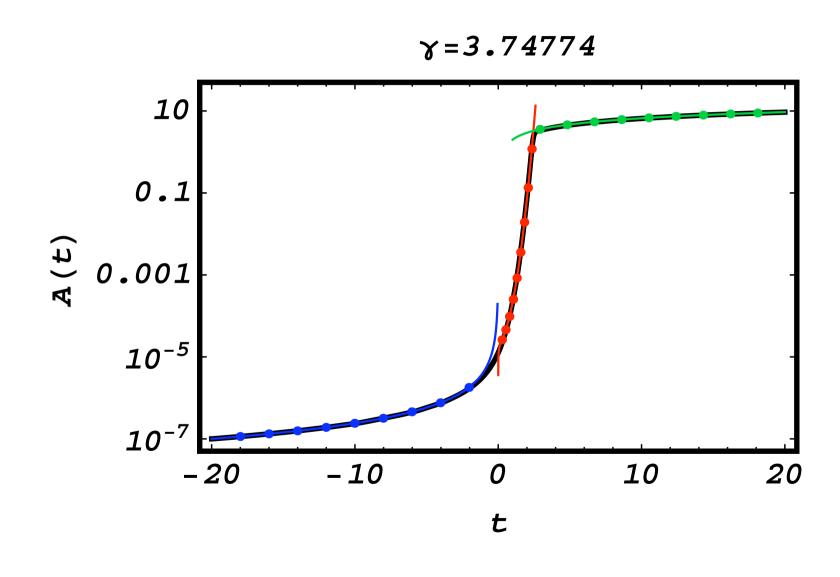
Nonlinear, Spatial Modulation

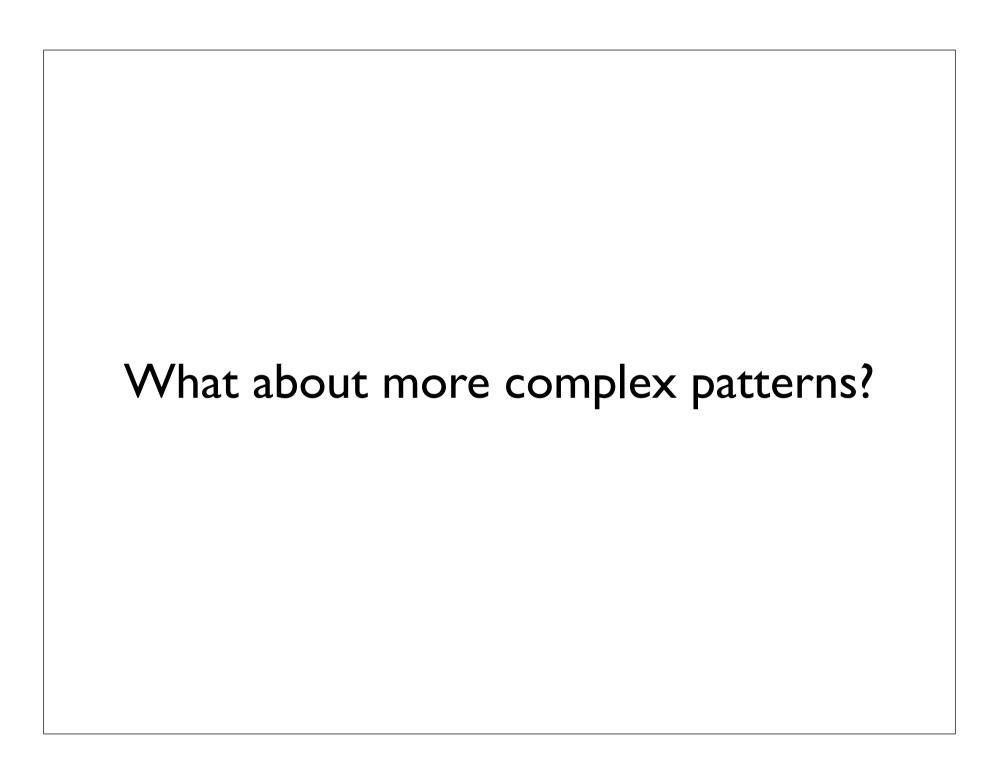
One of the most surprising new features...

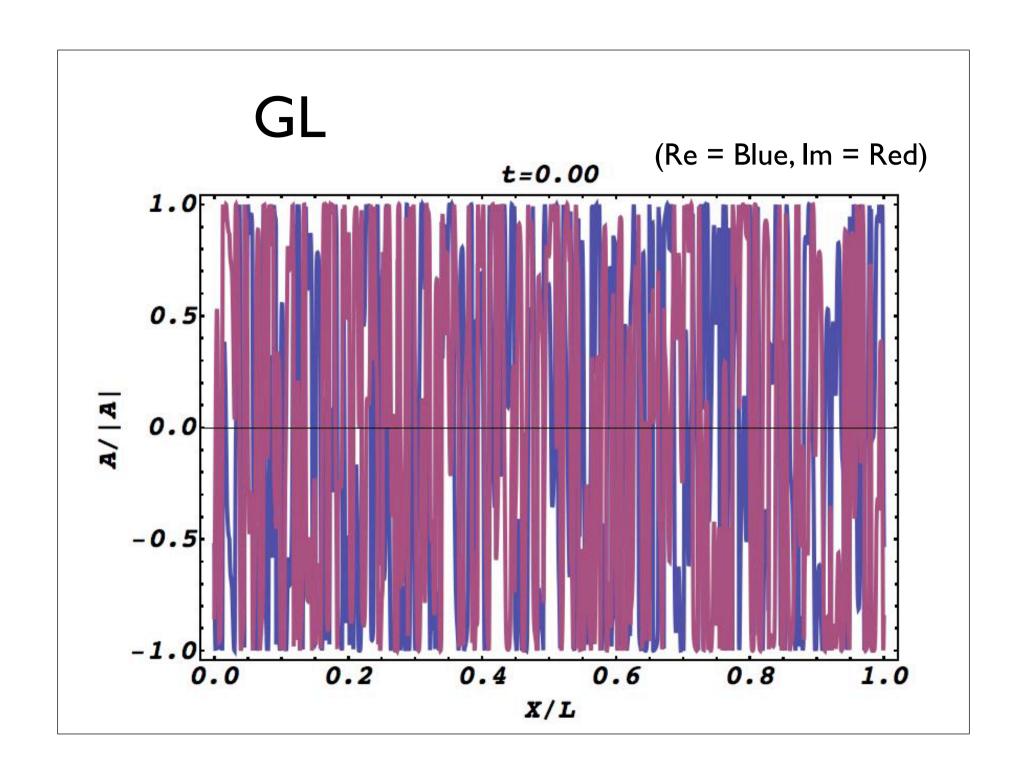
$$A o \sqrt{(\gamma + 2/3) \, au} \, e^{i\psi_0(X)} \, \operatorname{as} \, au o \infty$$

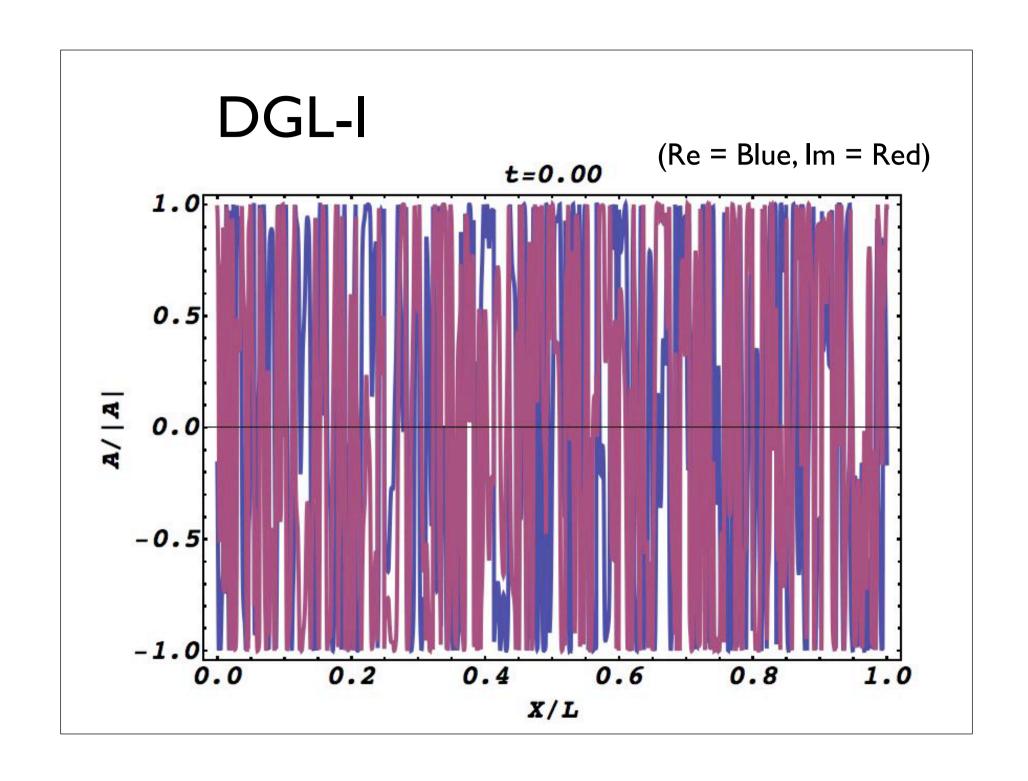
The "locked-in" phase function is strongly dependent on the initial conditions

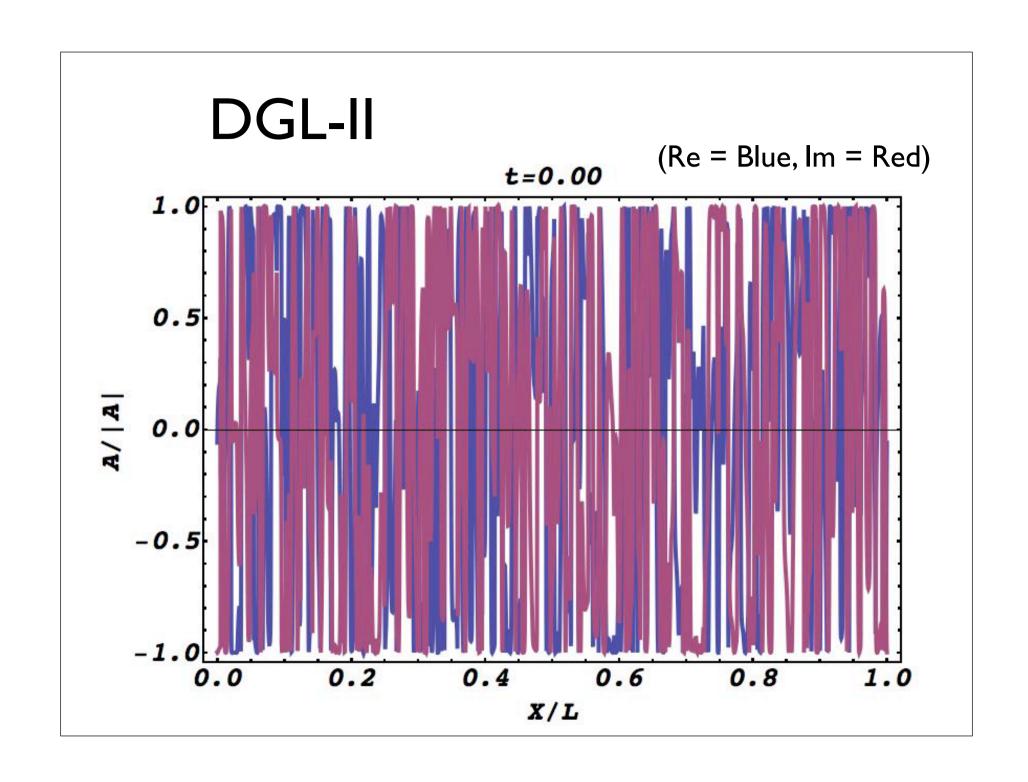
Purely Real, No Spatial Modulation

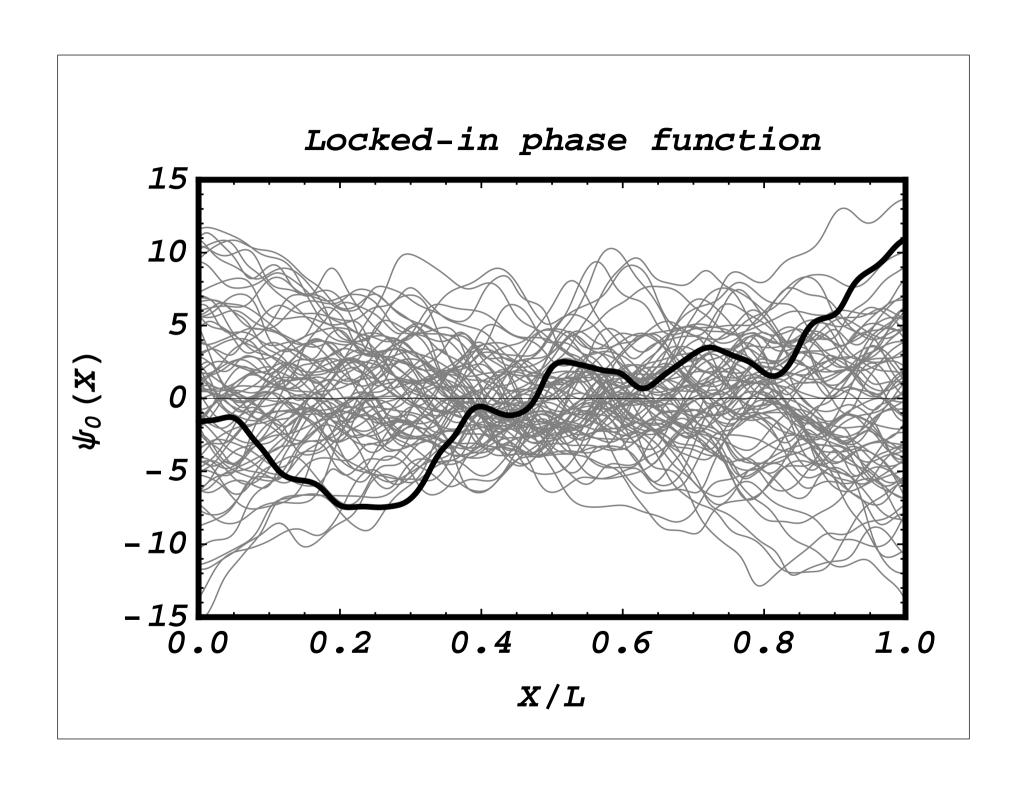


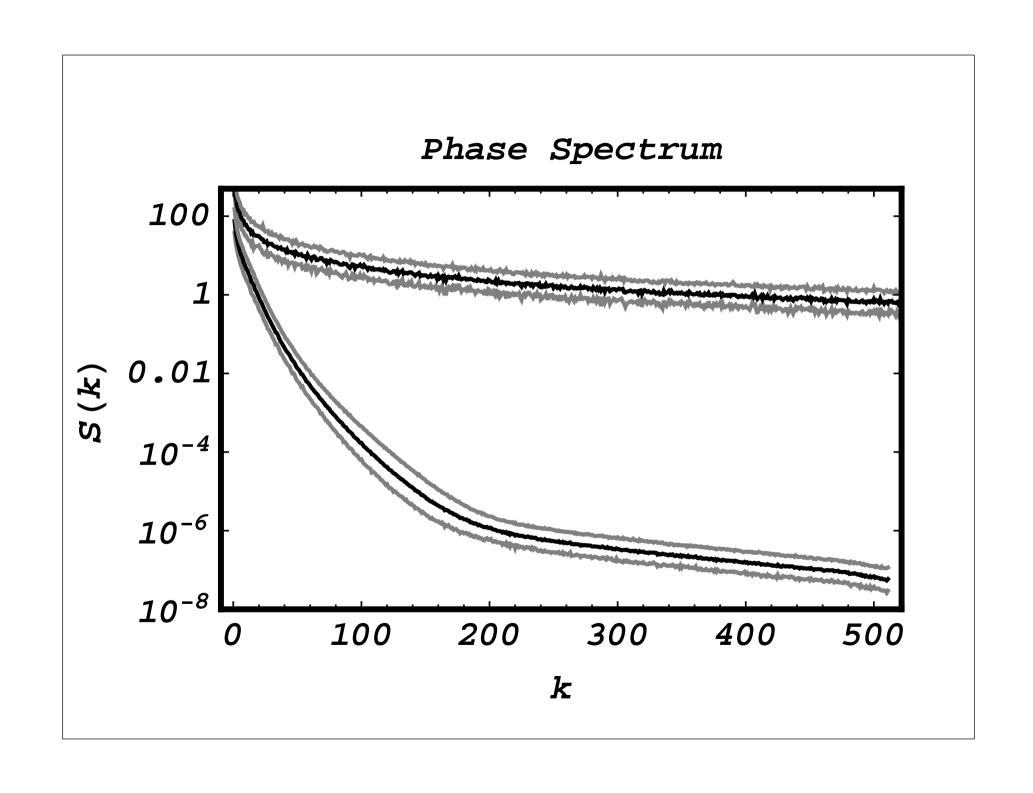




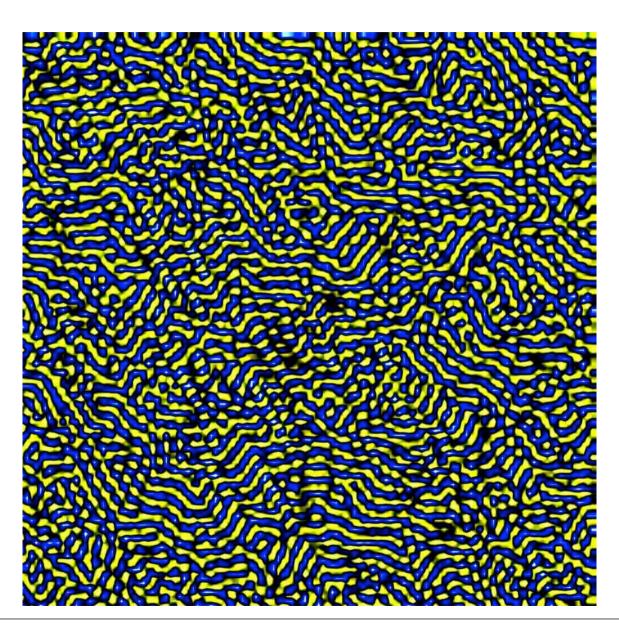








Patterns freeze in 2D as well



Some Thoughts...

Even strongly nonlinear systems often pass though a time when the growth rate of the instability is comparable to the rate of evolution of the background.

Conditions in this early stage can have dramatic consequences for the long-time dynamics of the system.

The pattern can even lock in broken-symmetry states that would otherwise be unstable.

We're hoping to set up an experiment with wax (or something) sometime soon.

I'm always looking for other systems to apply these basic methods to....